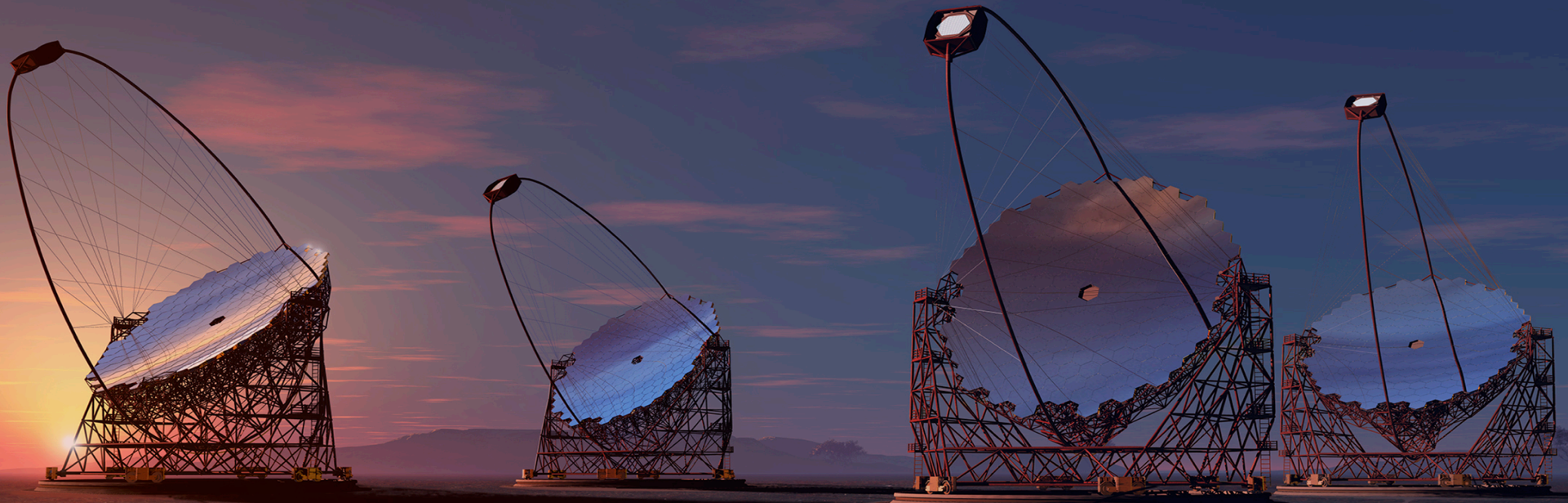




cherenkov  
telescope  
array

# Black Hole Astrophysics with the Cherenkov Telescope Array



Susumu Inoue (RIKEN)  
for the CTA Consortium

*S. Inoue*

# Outline



- 
- **The Cherenkov Telescope Array**
  - **Active Galactic Nuclei**
  - **Gamma-Ray Bursts**
  - **Microquasars, Other Transients**
  - **Extragalactic Survey**
  - **Summary**

Figures from “Science with CTA” document  
(to be published soon) unless noted otherwise



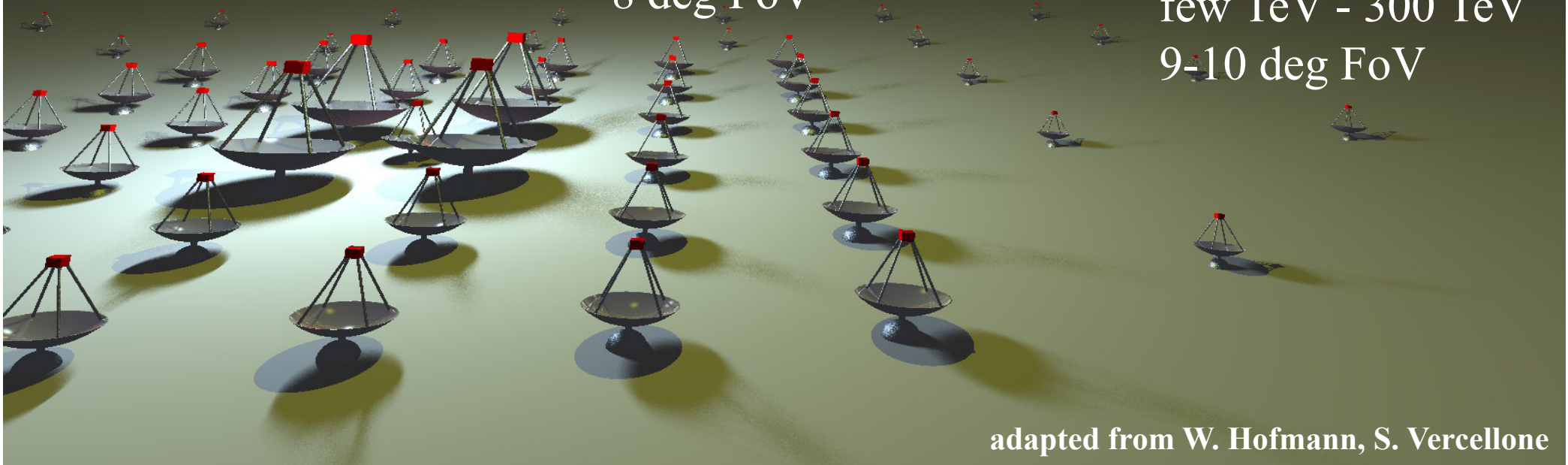
# The Cherenkov Telescope Array

- Array of state-of-the-art IACTs
- Combination of large-, middle-, small-sized telescopes for wide energy coverage
- 2 sites in N & S for all-sky coverage
- Open observatory for the community (some fraction of time allocated for Consortium Key Science Projects)

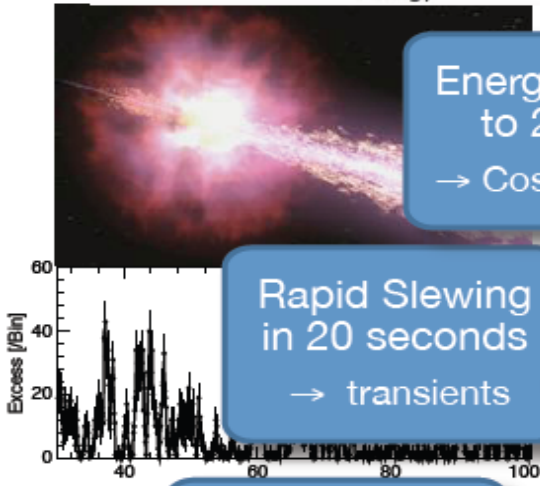
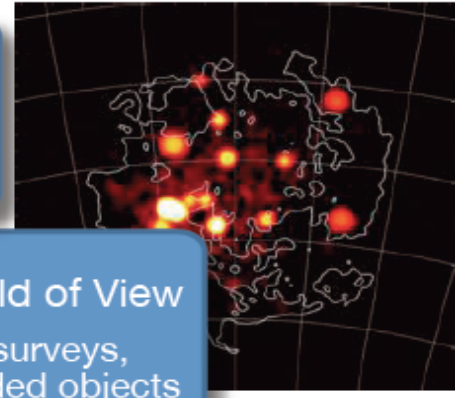
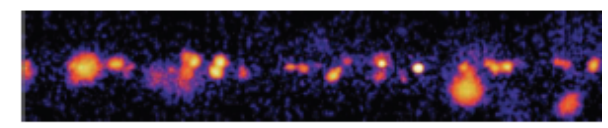
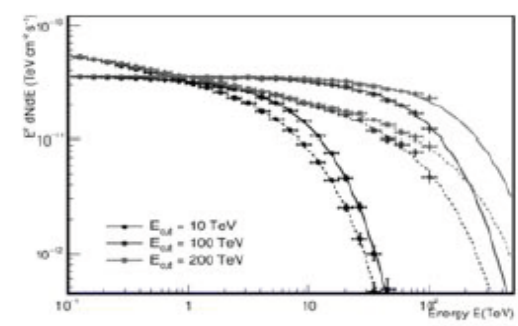
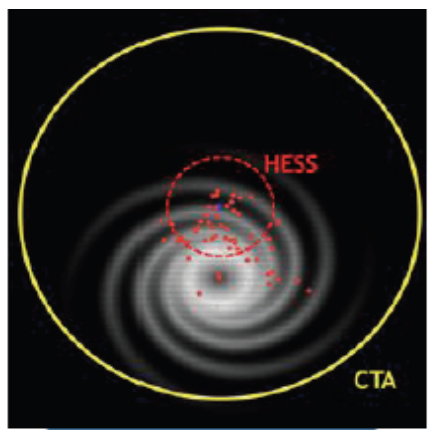
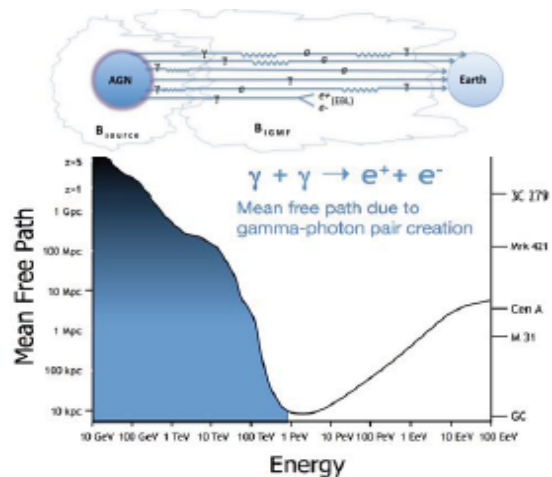
LSTs: 4 N, 4 S  
20 - 200 GeV  
4.5 deg FoV

MSTs: 25(+24)S, 15 N  
100 GeV - 10 TeV  
8 deg FoV

SSTs: 70 S  
few TeV - 300 TeV  
9-10 deg FoV



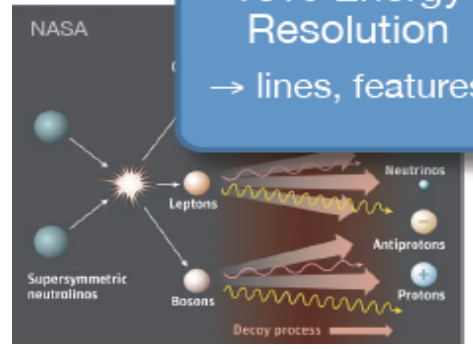
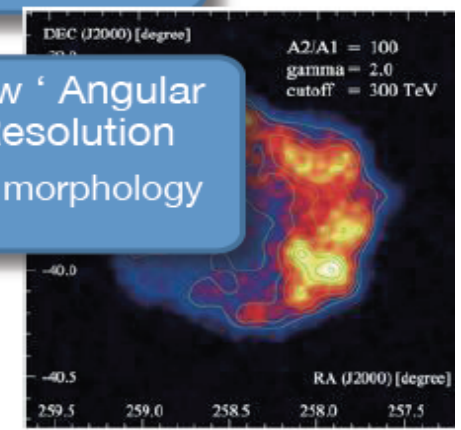
adapted from W. Hofmann, S. Vercellone



**cta**  
cherenkov telescope array

10% Energy Resolution  
→ lines, features

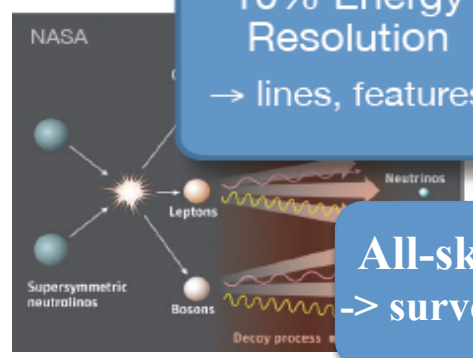
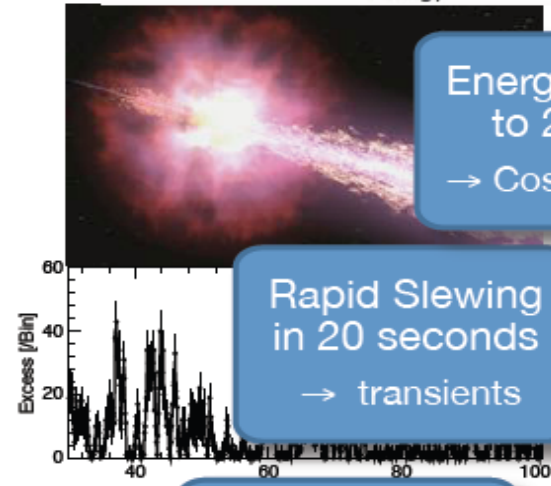
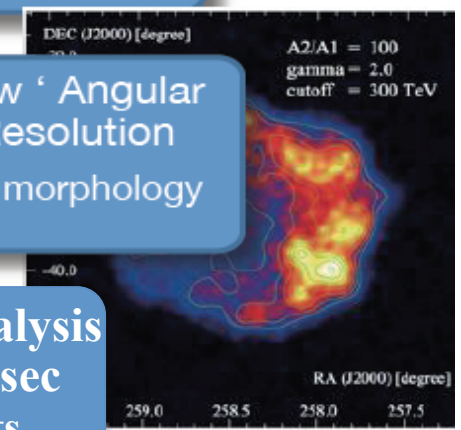
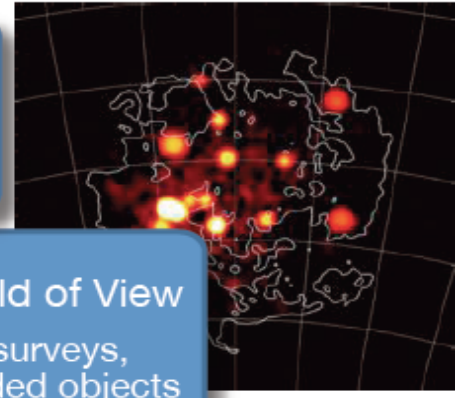
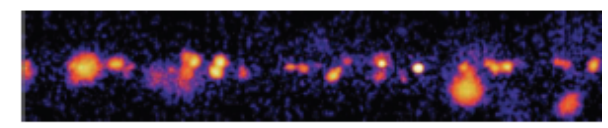
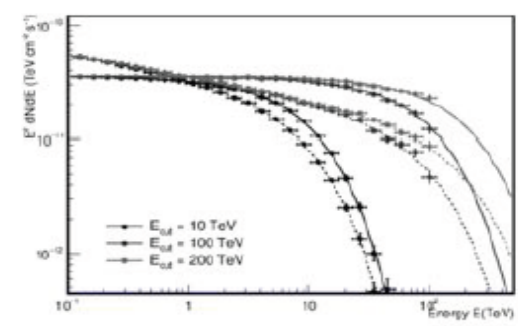
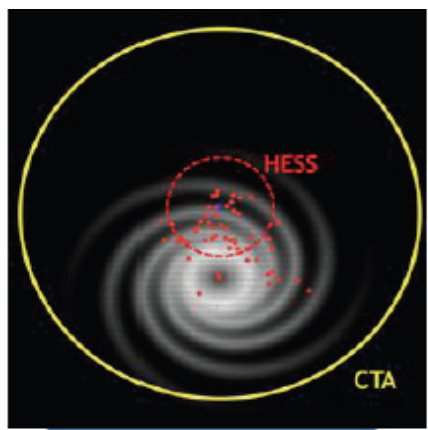
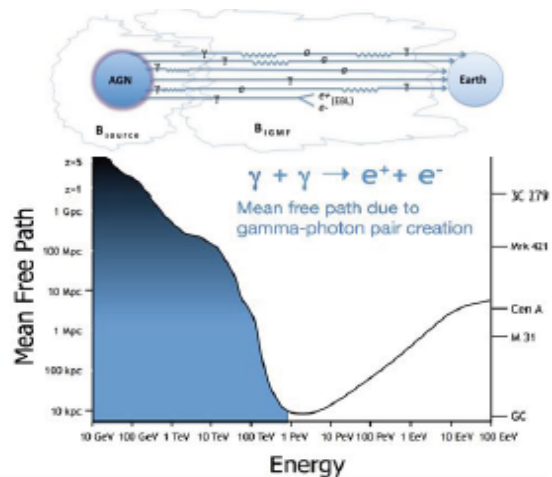
Few ' Angular Resolution  
→ morphology



# CTA vs current IACTs

from W. Hofmann

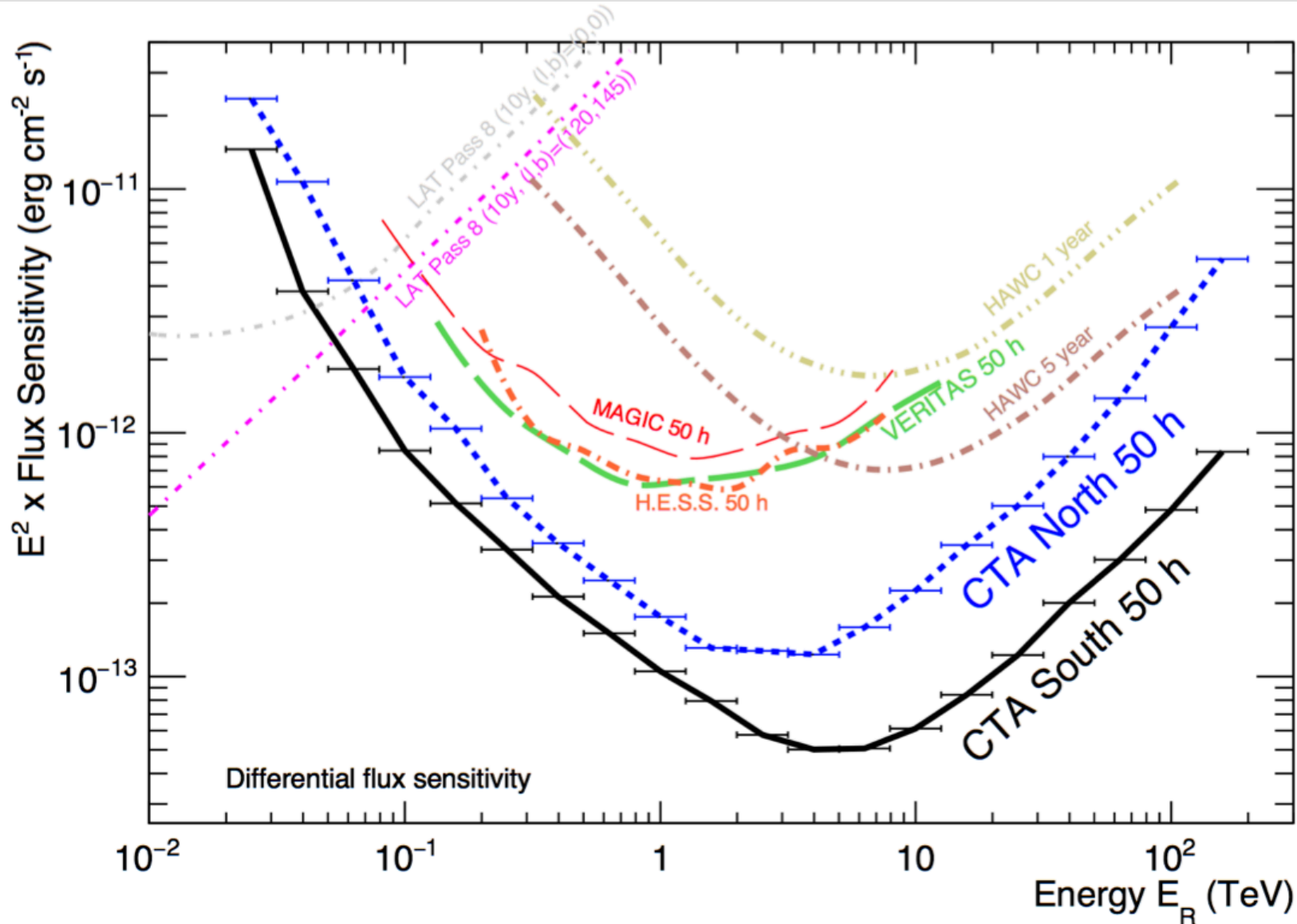




# CTA vs current IACTs

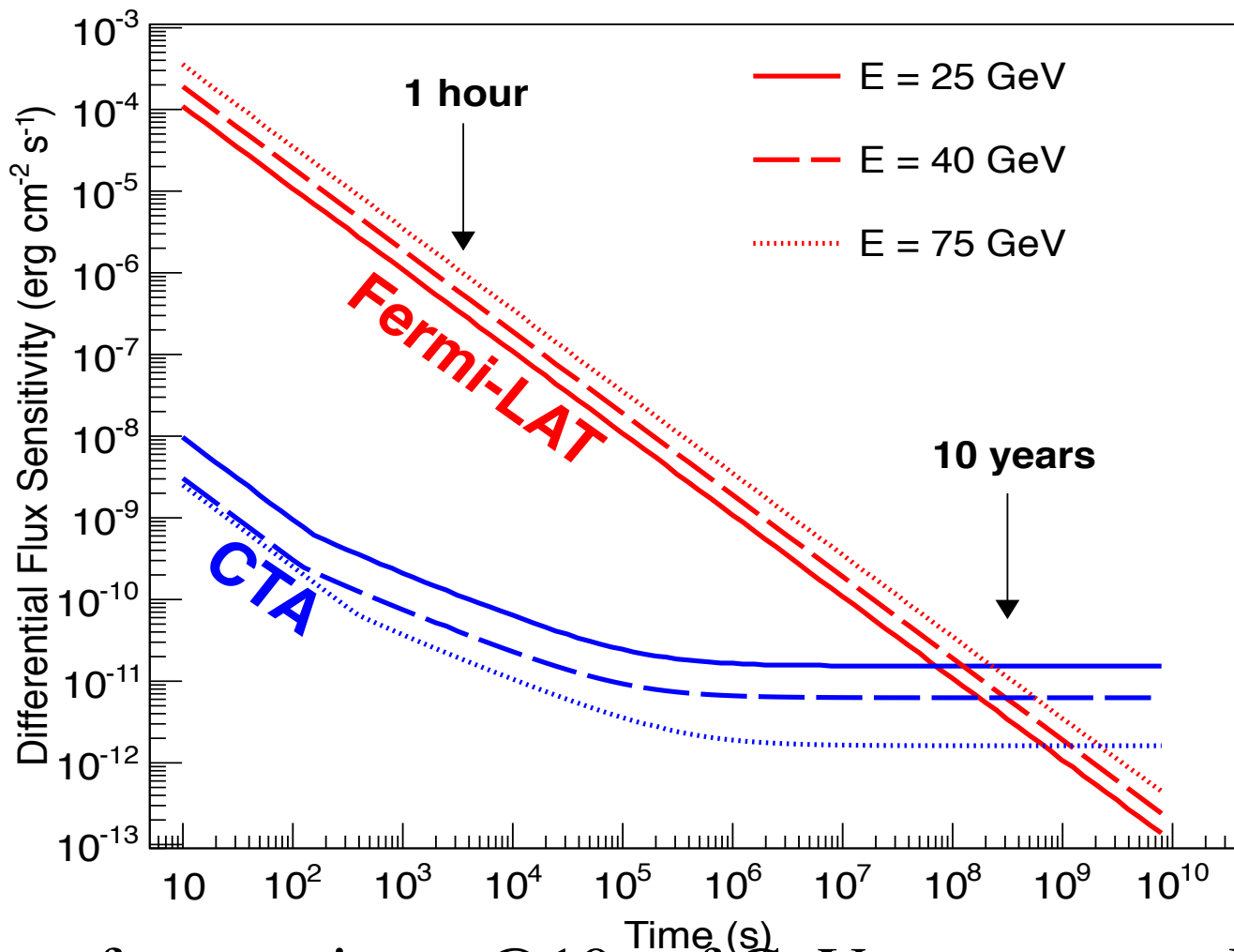
adapted from W. Hofmann

# CTA Sensitivity (steady sources)





# CTA Sensitivity vs Fermi



Pros:

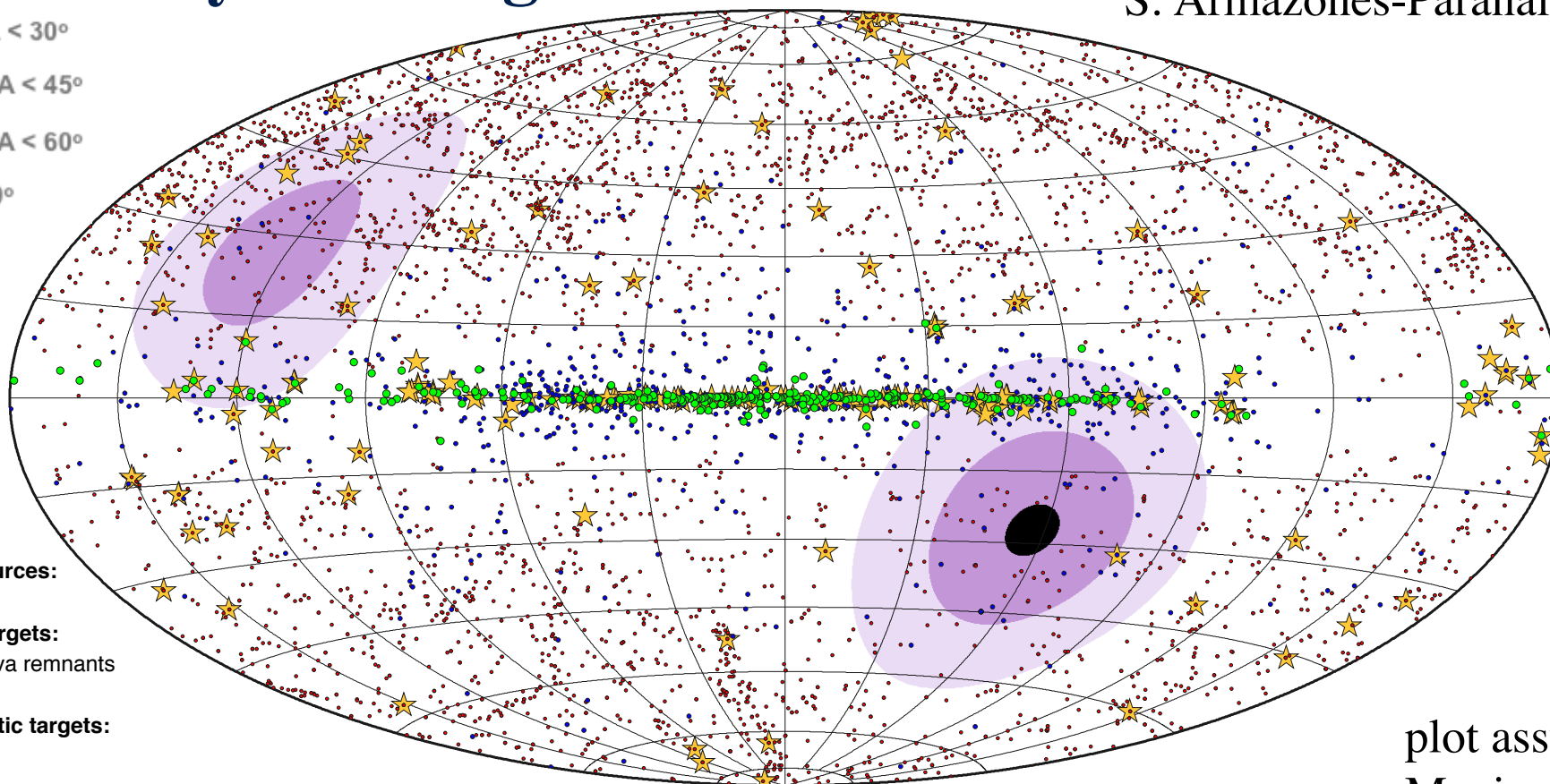
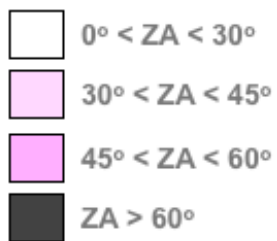
Big advantage for transients @ 10s of GeV  
effective area  $\sim 10^4 \times \text{LAT}@30\text{GeV}$

Cons:

Limited FoV  
Limited duty cycle

# All Sky Coverage

N: La Palma, Canary Is.  
S: Armazones-Paranal, Chile



**Known sources:**

★ TeVCat

**Galactic targets:**

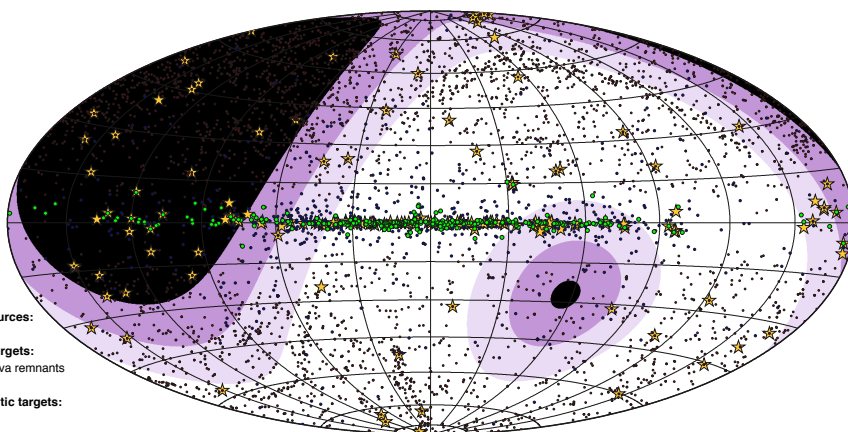
● Supernova remnants

● Pulsars

**Extragalactic targets:**

● Blazars

plot assumes  
Mexico for N



**Known sources:**

★ TeVCat

**Galactic targets:**

● Supernova remnants

● Pulsars

**Extragalactic targets:**

● Blazars

**Known sources:**

★ TeVCat

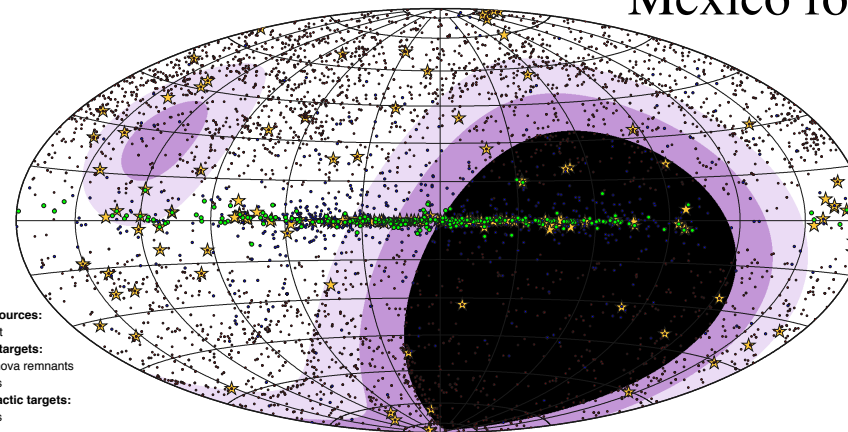
**Galactic targets:**

● Supernova remnants

● Pulsars

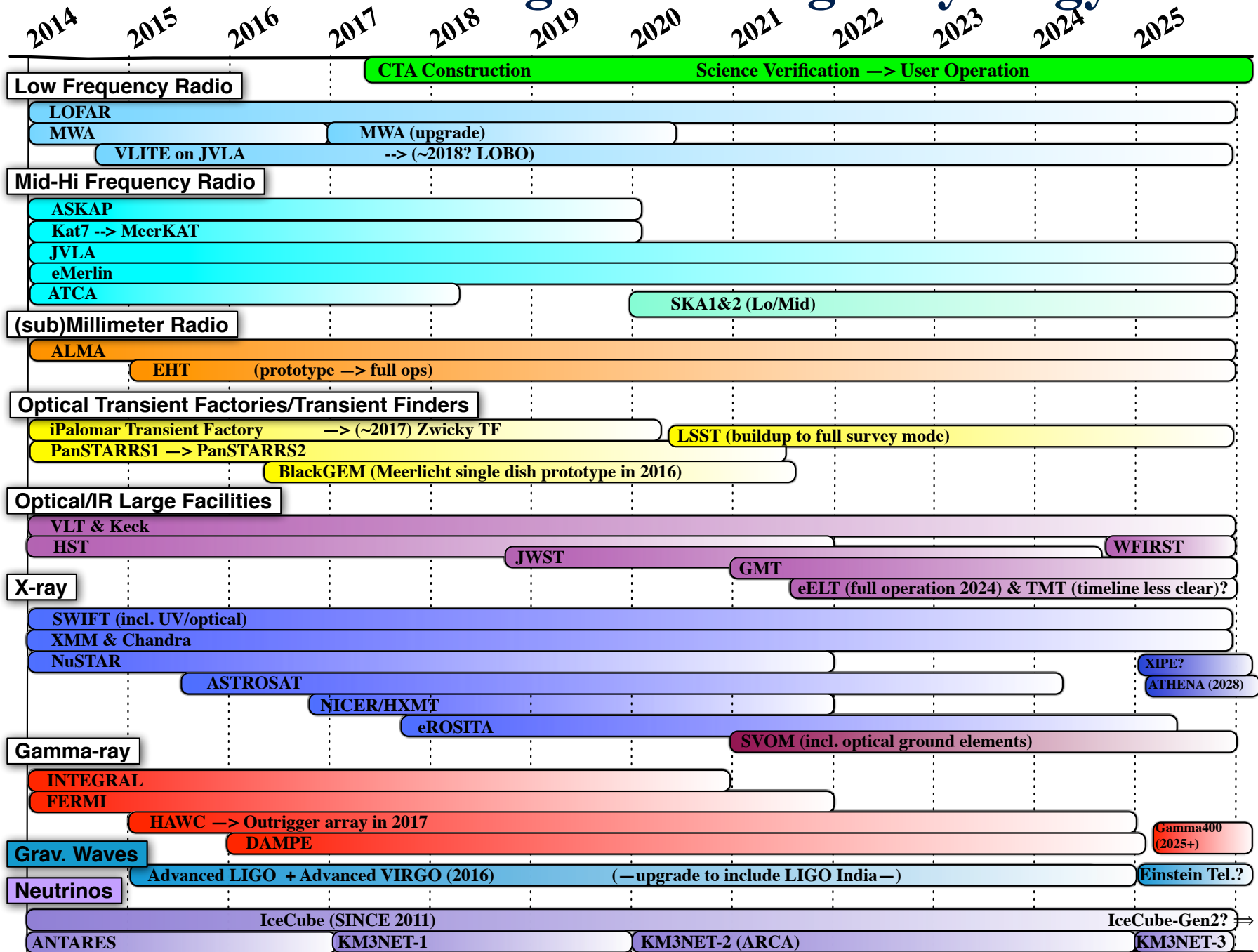
**Extragalactic targets:**

● Blazars





# Multi-wavelength/messenger Synergy



# CTA Key Science Projects

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- **Dark Matter Programme**
- **Galactic Centre**
- **Galactic Plane Survey**
- **Large Magellanic Cloud Survey**
- **Extragalactic Survey**
- **Transients**
- **Cosmic Ray PeVatrons**
- **Star Forming Systems**
- **Active Galactic Nuclei**
- **Clusters of Galaxies**
- **Non-Gamma-ray Science**

Details in document  
“Science with CTA”  
to be published on arXiv



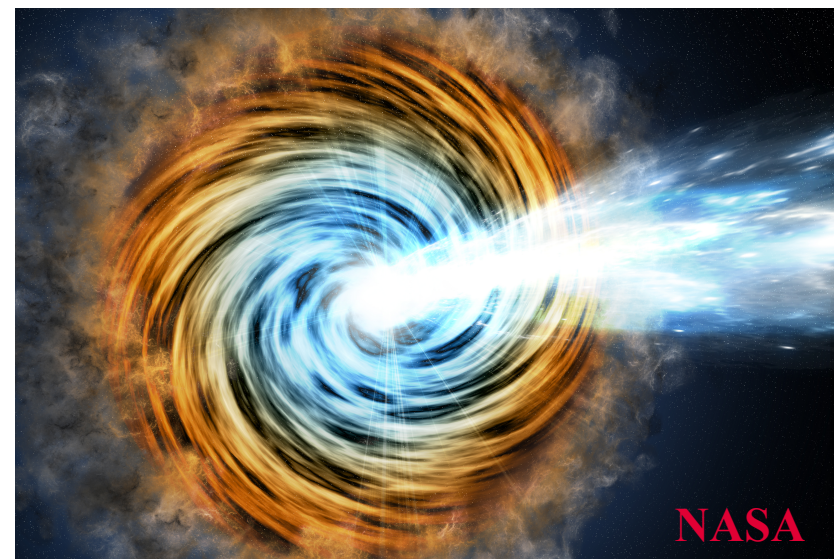
# Active Galactic Nuclei



- **Physics of relativistic jets from supermassive BHs**
  - Mechanisms: emission, particle acceleration, energy dissipation, jet formation
  - Demography: origin of diversity, search for new classes
- **Tools to probe the Universe**
  - Extragalactic background light (star formation history, etc)
  - Intergalactic magnetic fields
- **Tests of UHECR origin, fundamental physics**

Search for signatures of:

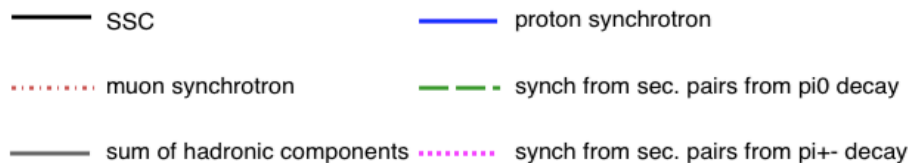
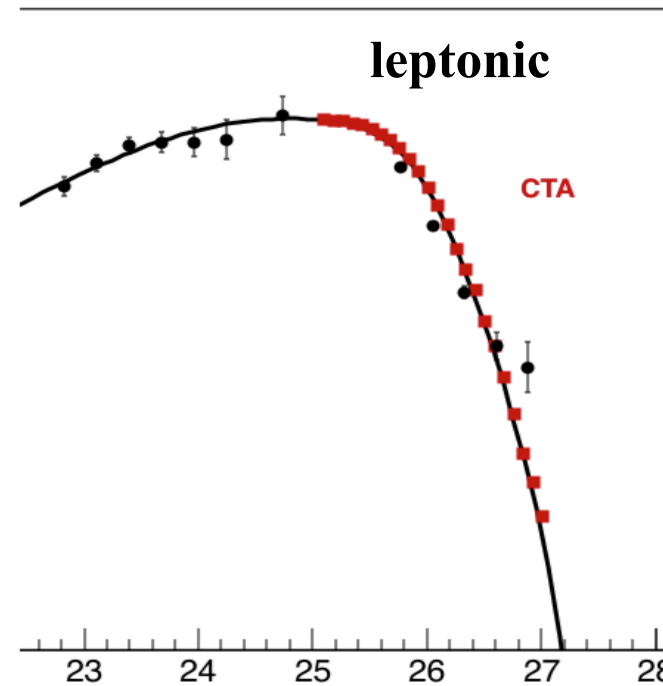
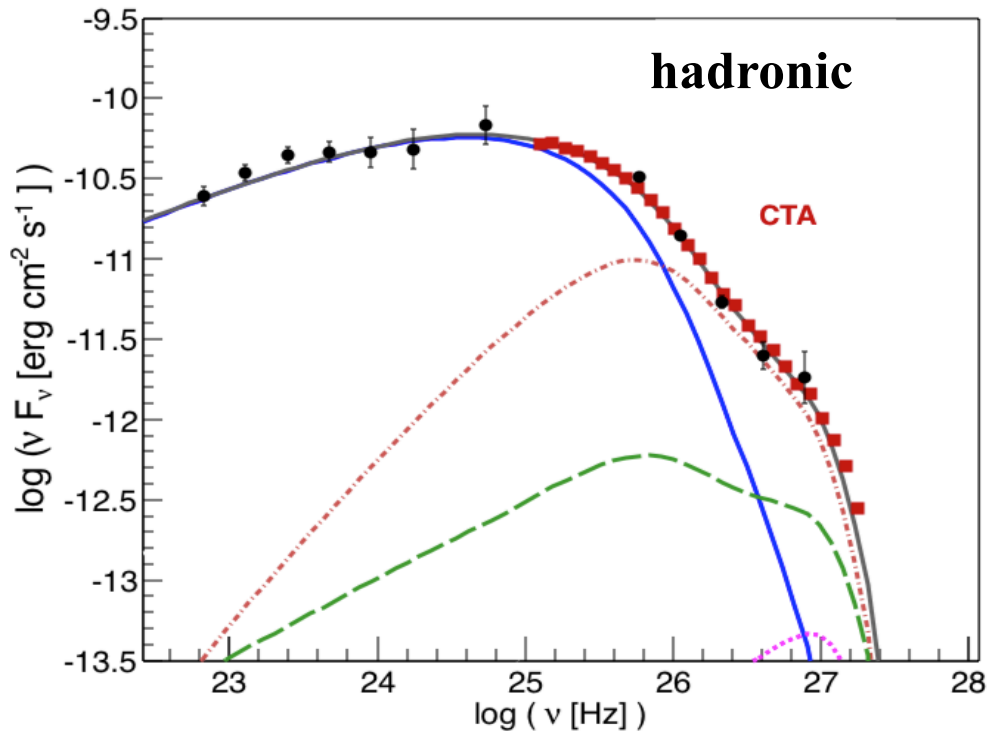
  - Accelerated hadrons
  - Lorentz invariance violation, Axion-like particles



# Active Galactic Nuclei



- Clarify physics of emission, particle acceleration
- Test UHECR origin high S/N spectra + variability



PKS 2155-304

■ H.E.S.S. data

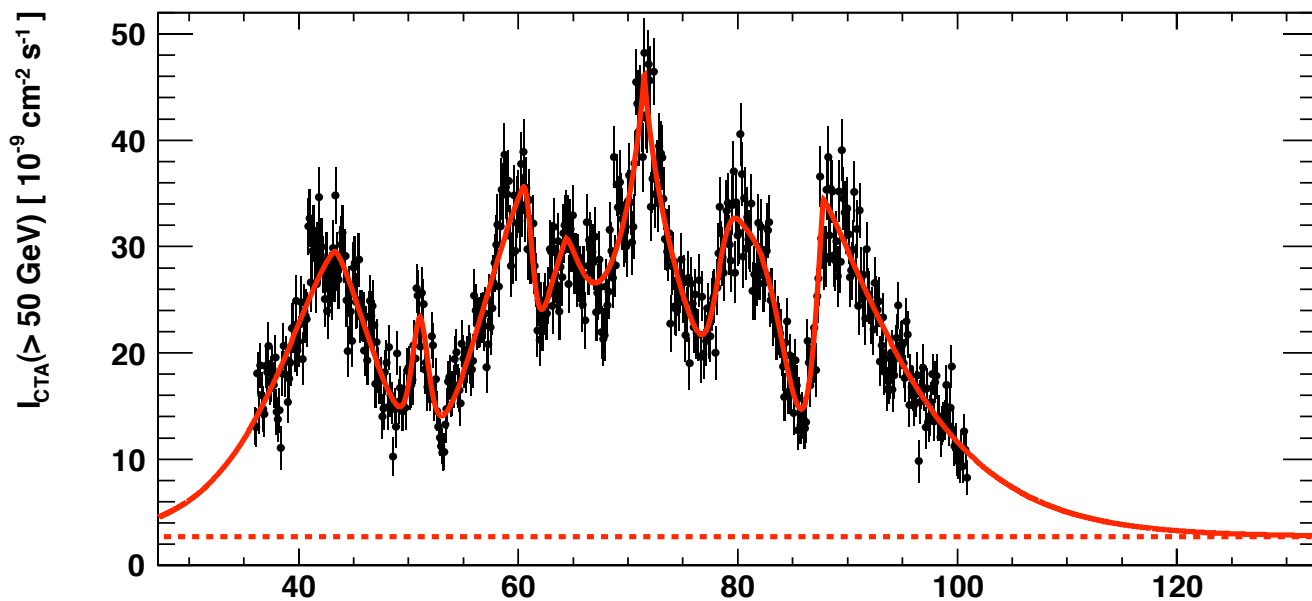
■ 33hr CTA simulation



# Active Galactic Nuclei

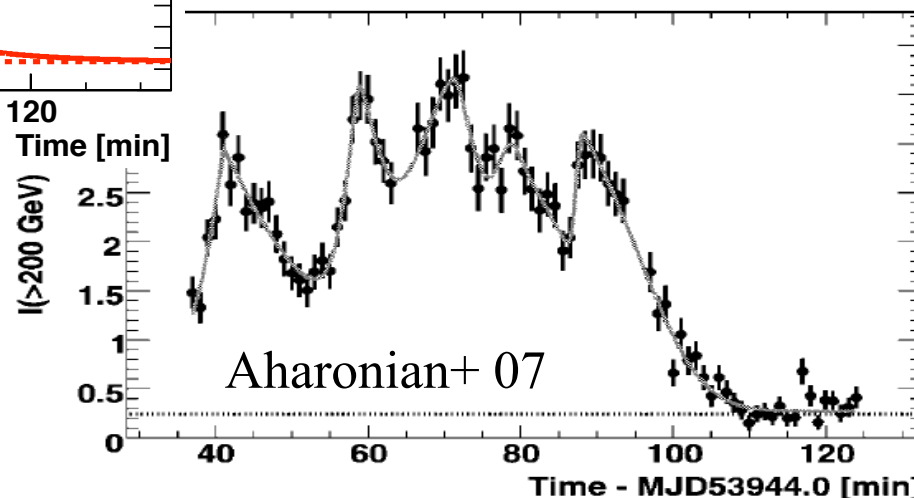


- Clarify physics of emission, particle accel., energy dissipation
- Test LIV high S/N light curves



CTA simulation  
 $\Delta t_{\text{min}} \sim 25 \text{ sec}$

**PKS 2155-304**  
HESS obs.  
 $\Delta t_{\text{min}} \sim 3 \text{ min}$

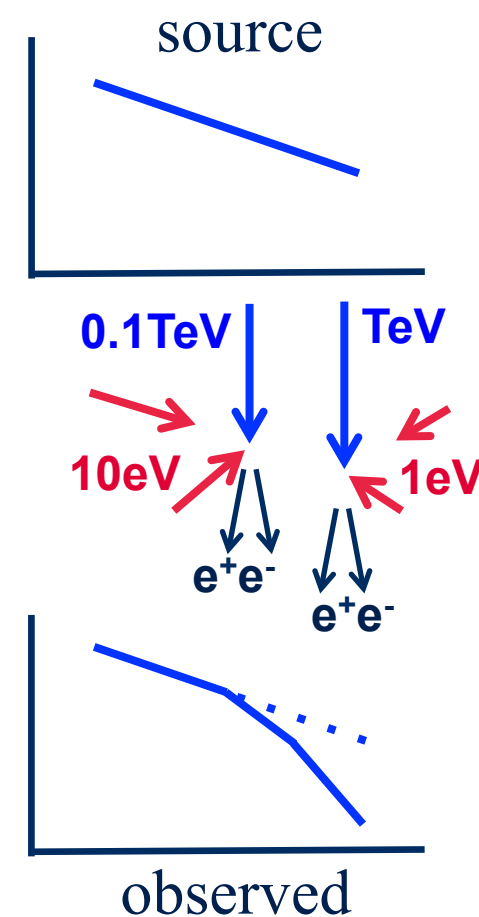
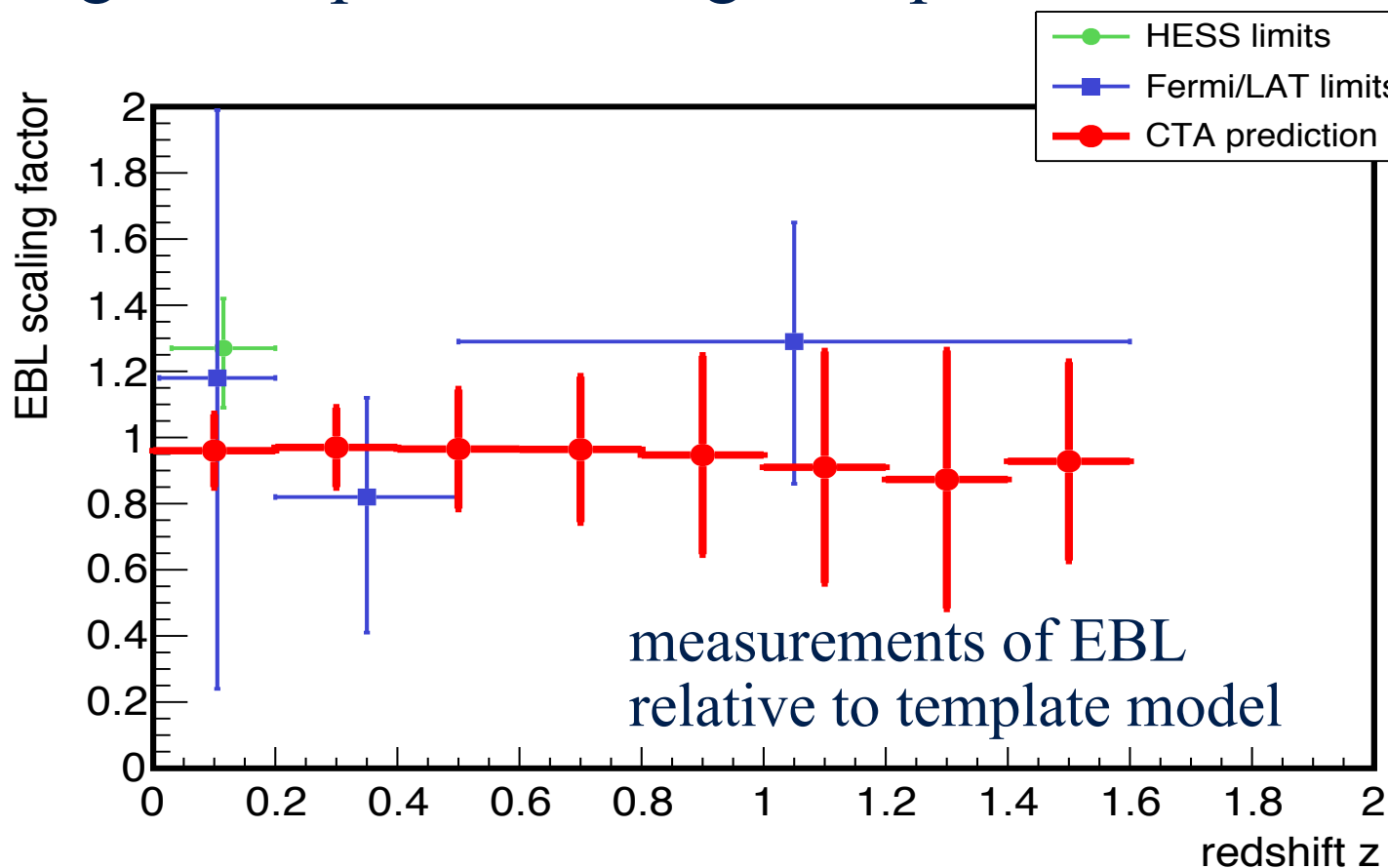


Aharonian+ 07

# Active Galactic Nuclei



- Probe evolution of optical/IR EBL via  $\gamma\gamma$  spectral attenuation
- global view of star/galaxy formation & evolution
- high S/N spectra for large sample with different  $z$



# Gamma-Ray Bursts



- **Physics of GRBs**
  - Prompt: mechanism, jet properties, central engine (NS or BH?)
  - Early afterglow: mechanism (plateau phase), particle acceleration, B field generation
- **Tools to probe the Universe**
  - Extragalactic background light (deeper than AGN)
  - Intergalactic magnetic fields
- **Tests of UHECR origin, fundamental physics**

Search for signatures of:

  - Accelerated hadrons
  - Lorentz invariance violation

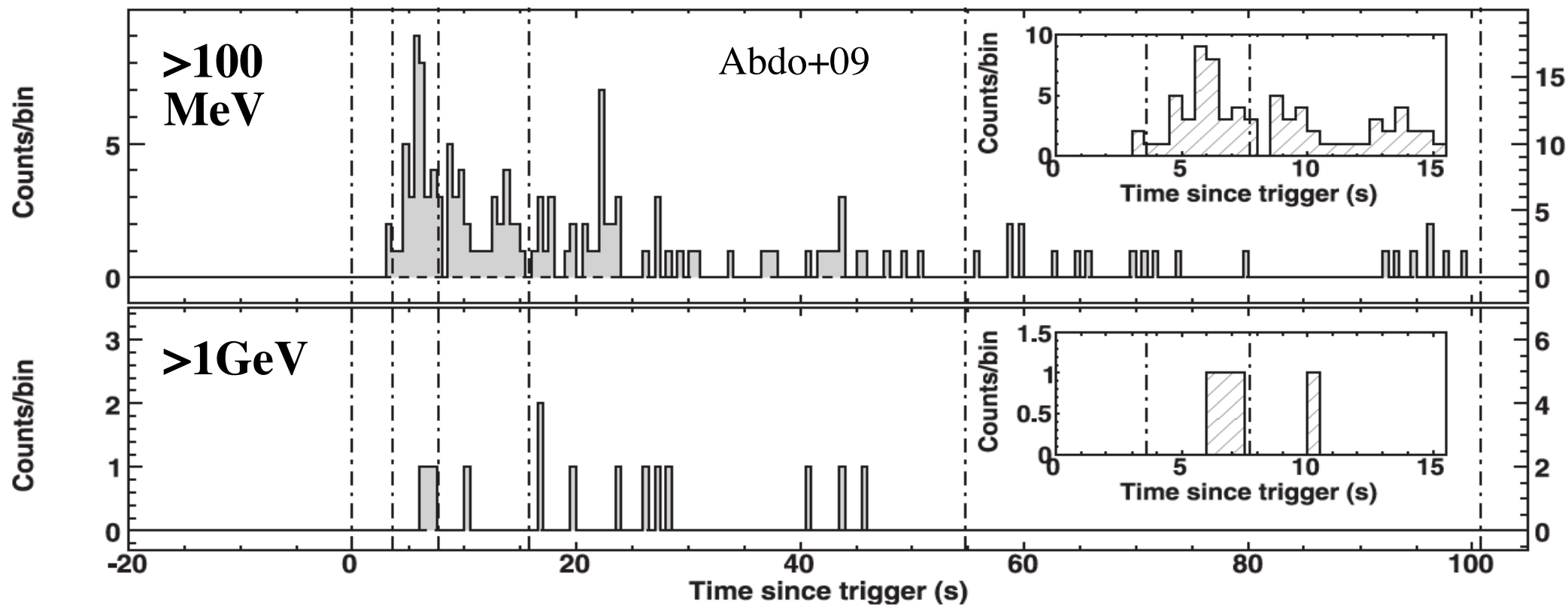




# Gamma-Ray Bursts



## Fermi-LAT observations GRB 080916C

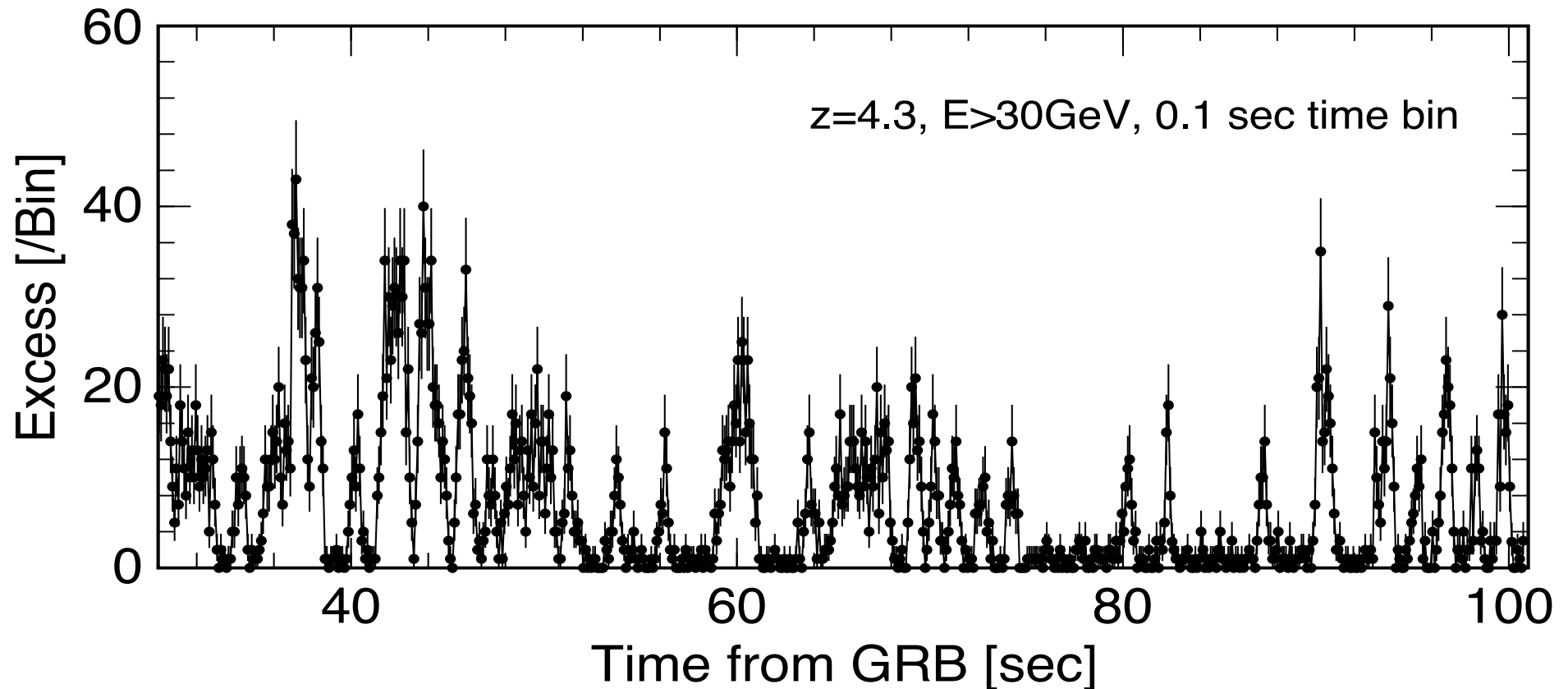


# Gamma-Ray Bursts



- Clarify physics of emission
  - Test UHECR origin, LIV
- high S/N, E-dependent light curves

CTA simulation  
**GRB 080916C**



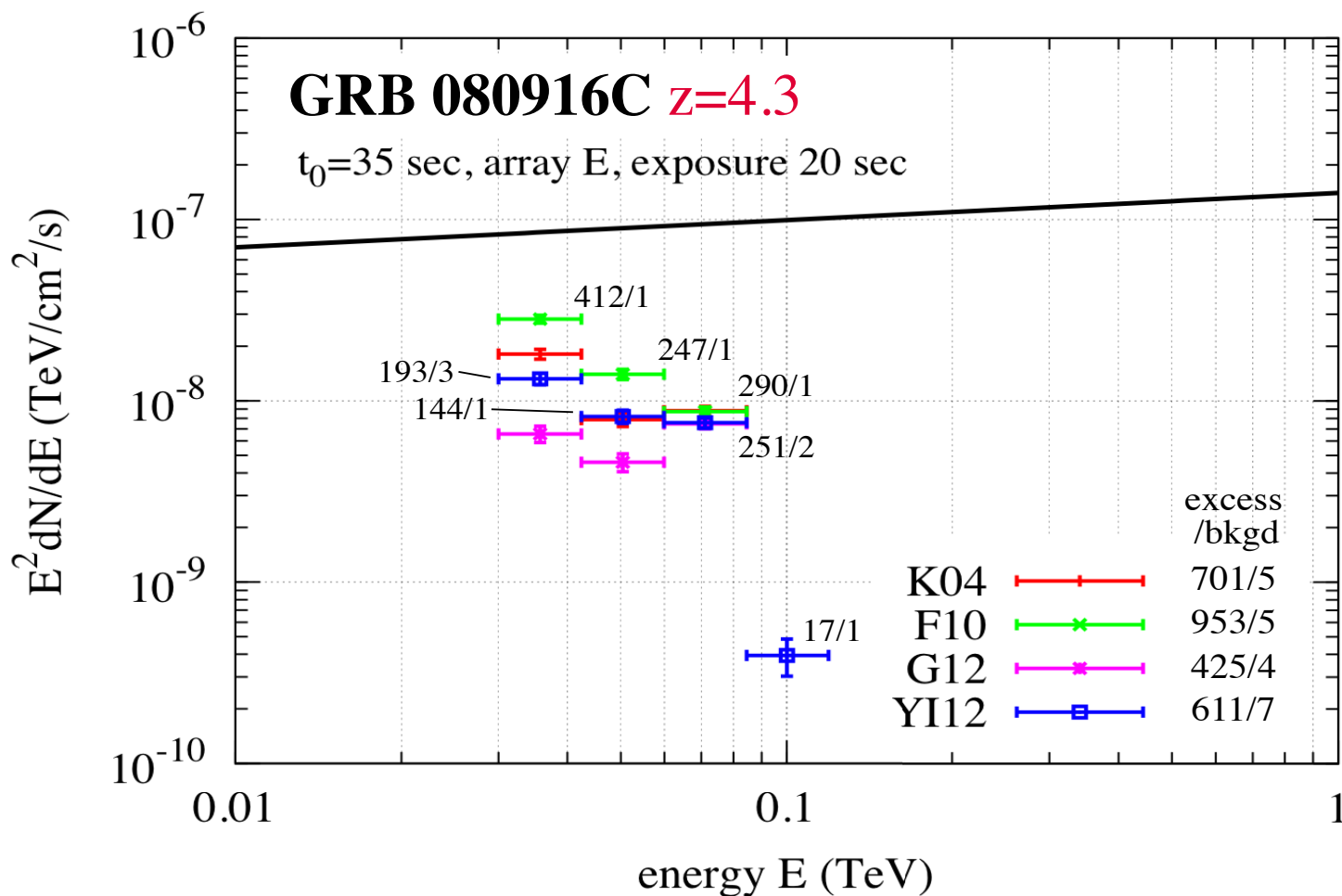
# Gamma-Ray Bursts



- Probe high-z UV EBL via  $\gamma\gamma$  spectral attenuation

high S/N spectra

CTA simulation





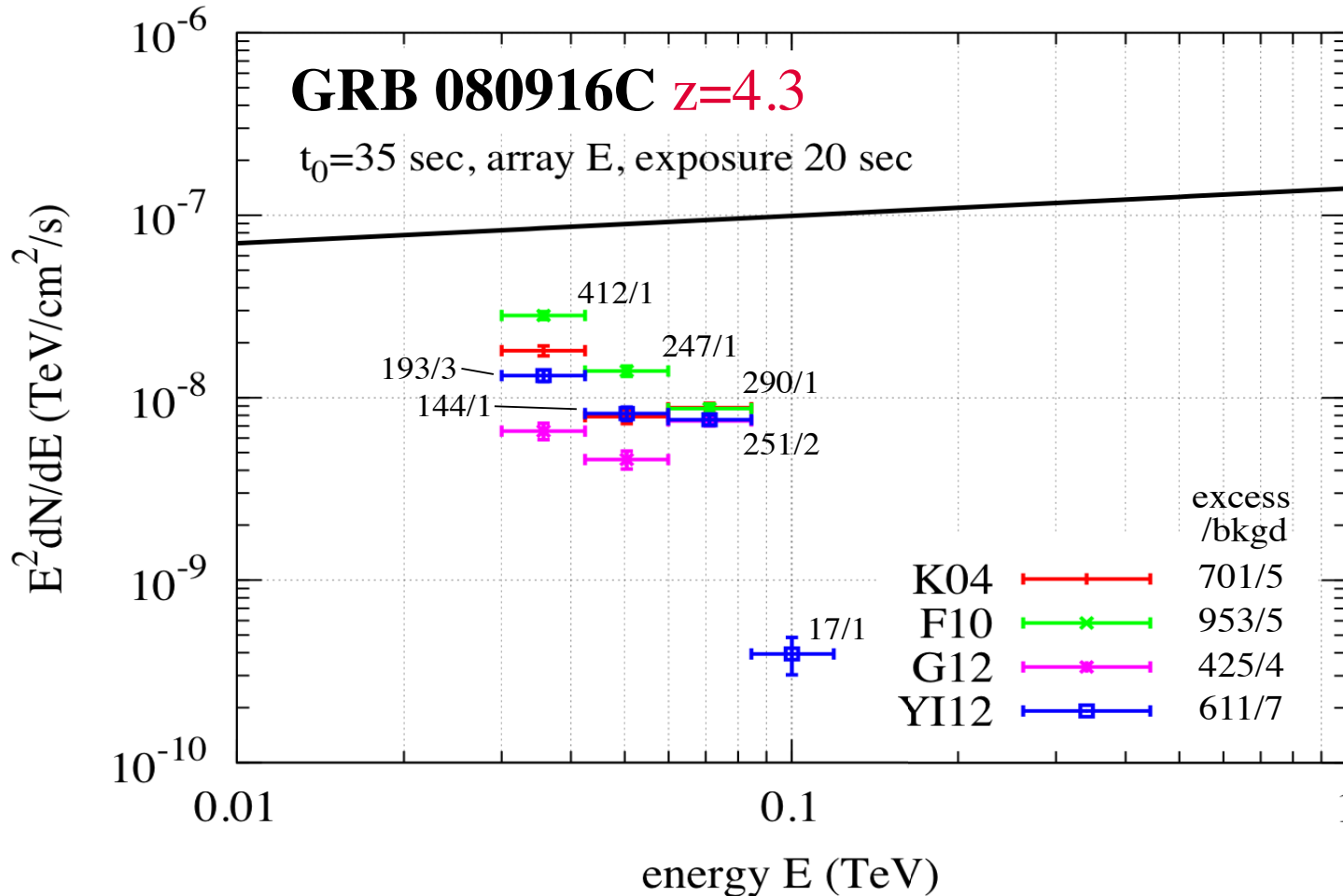
# Gamma-Ray Bursts



## - Probe high-z UV EBL via $\gamma\gamma$ spectral attenuation

high S/N spectra

CTA simulation



expected total  
detection rate:  
~1-2 /yr/site  
~<15% prompt  
rest afterglow

Kakuwa+ 12

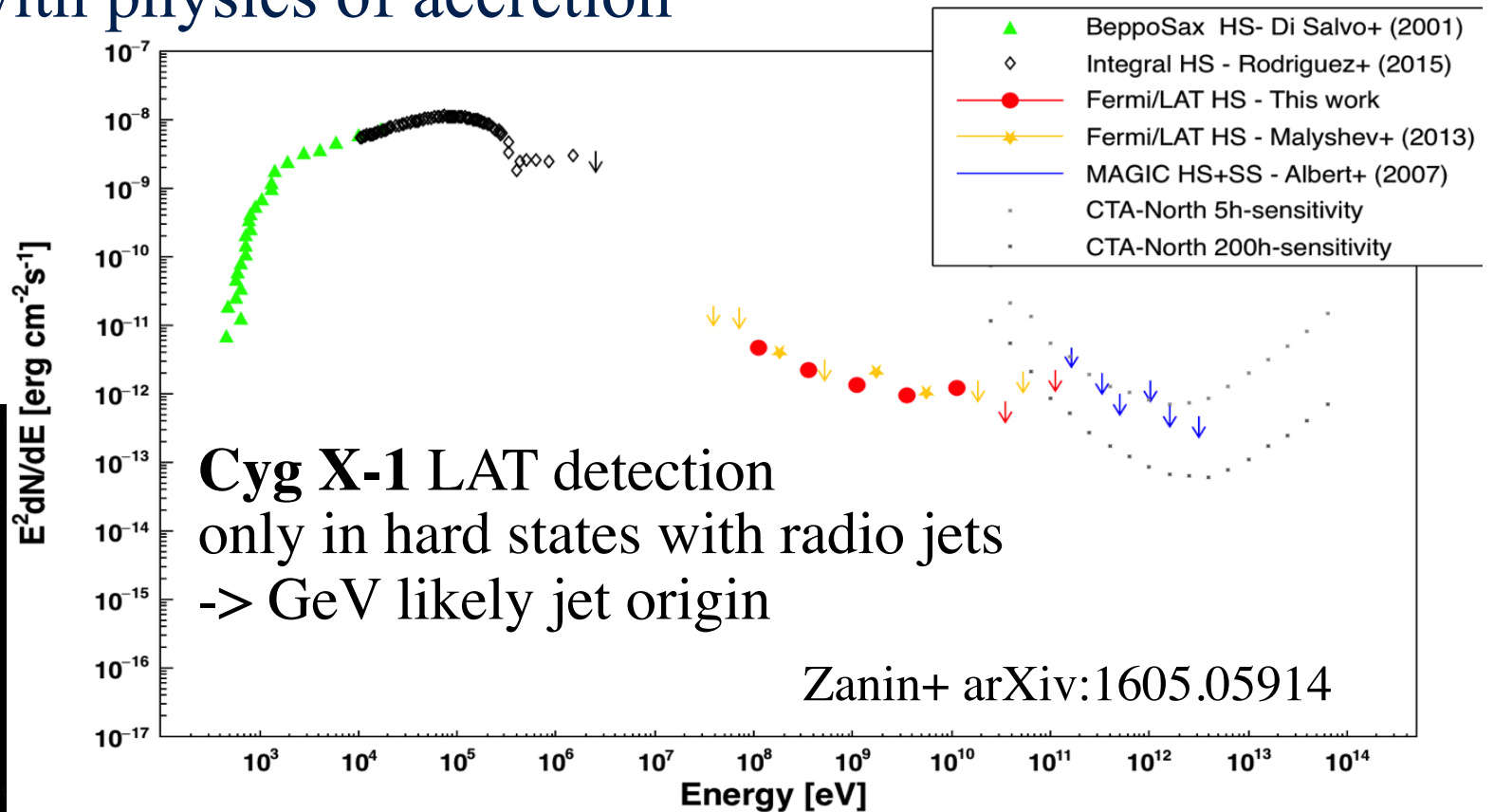
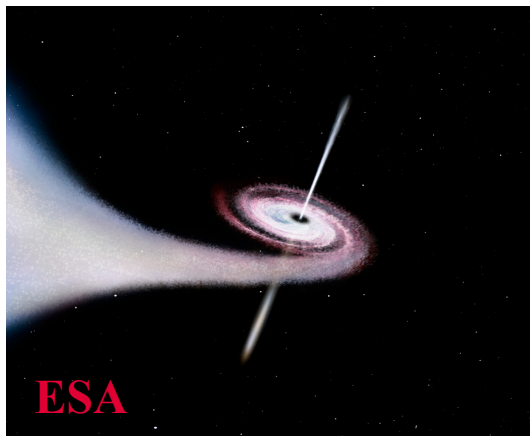
Gilmore+ 13

SI+ 13

# Microquasars



- **Physics of relativistic jets from stellar mass BHs**
  - Mechanisms: emission, particle acceleration, energy dissipation, jet formation
  - + connection with physics of accretion



# Other Transients



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## Follow-up of Alerts:

- **High-Energy Neutrinos**
- **Gravitational Waves**
- **X-ray/Optical Transients**
  - Tidal Disruption Events
  - Supernova Shock Breakout Events
  - ...
- **Radio Transients**
  - Fast Radio Bursts
  - ...

possibly with  
subarrays

## Serendipitous Transients

via real time analysis, alerts with  $\sim 30$  sec latency

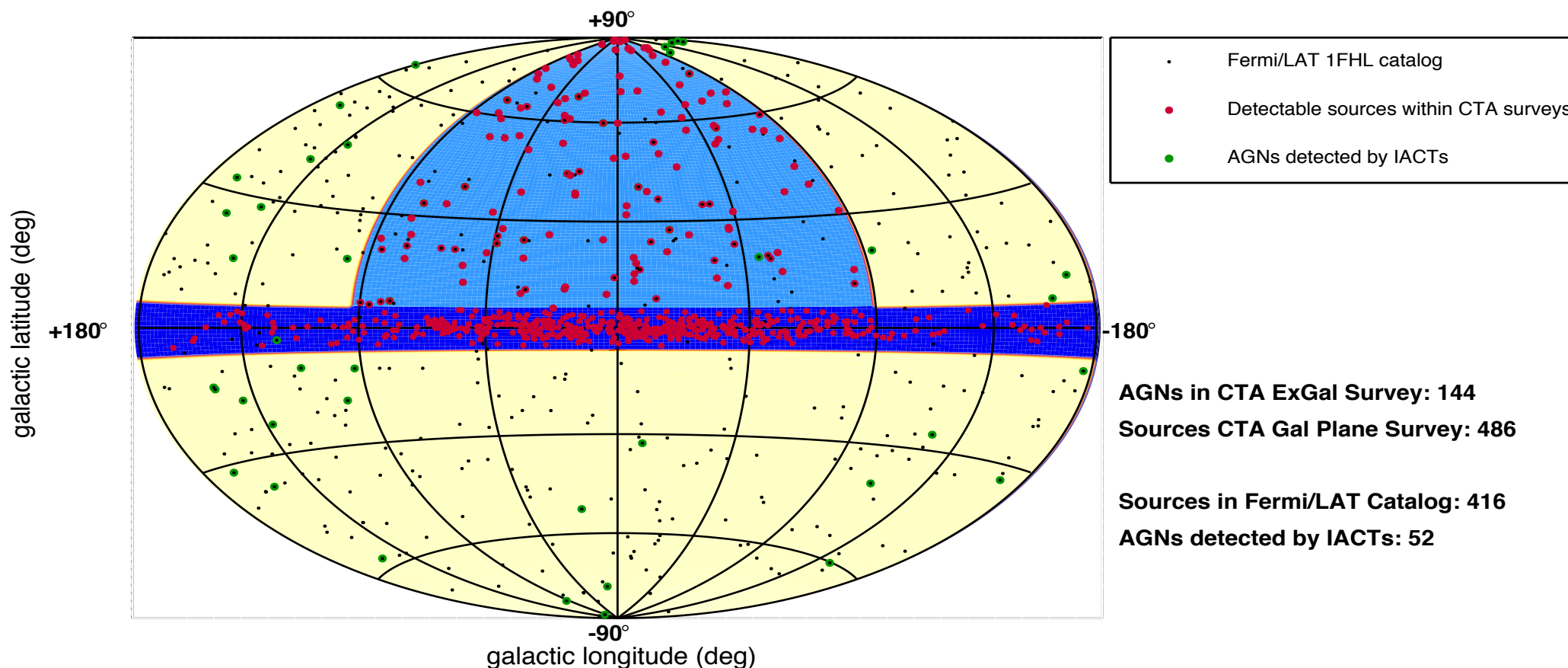
-> CTA as transient factory



# Extragalactic Survey



Unbiased survey for 1/4 sky ( $\sim 10^4$  deg<sup>2</sup>) to flux limit  $\sim 5$ mCrab

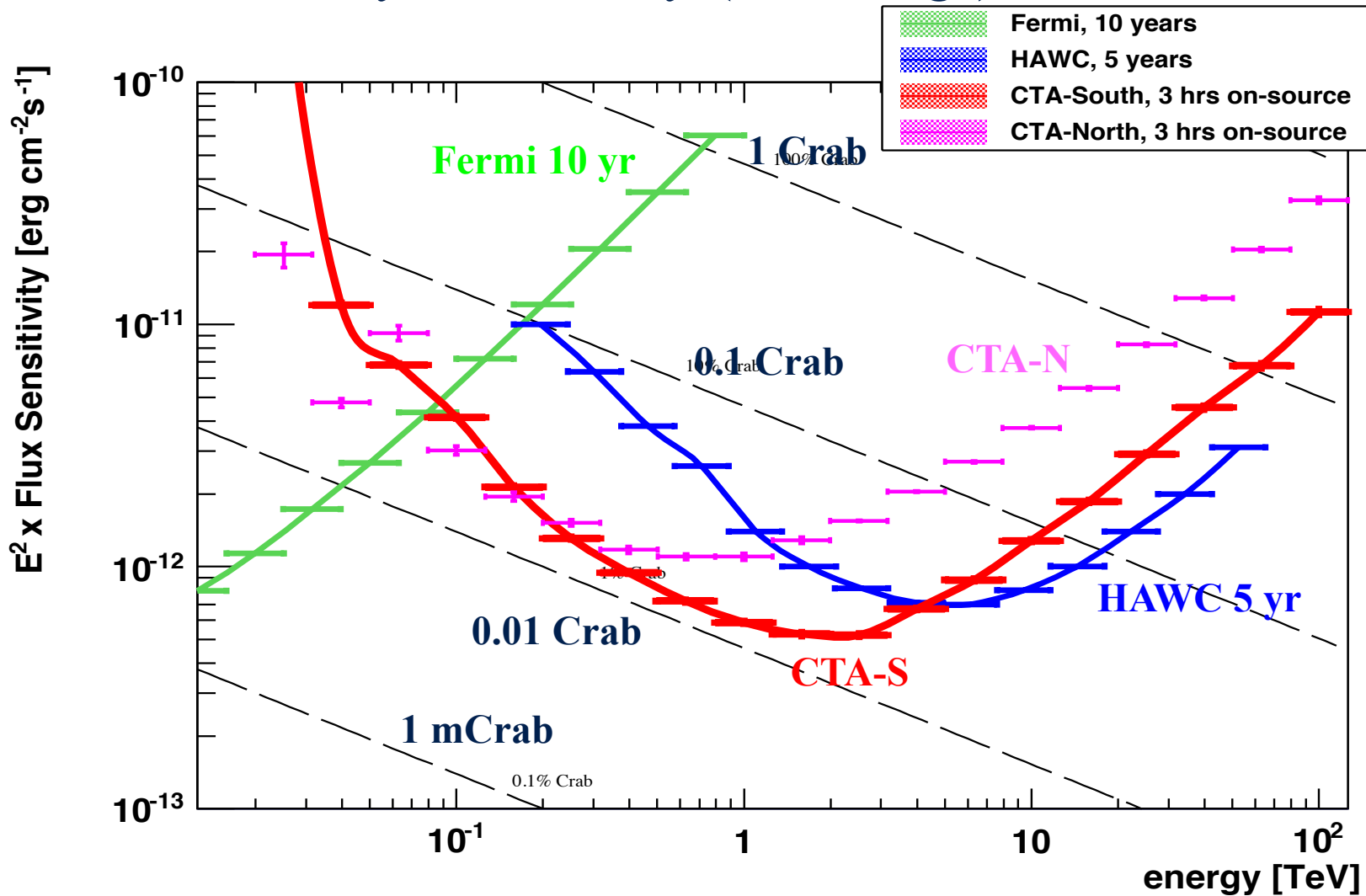


covers Virgo, Coma, Cen A, North Fermi Bubble

# Extragalactic Survey



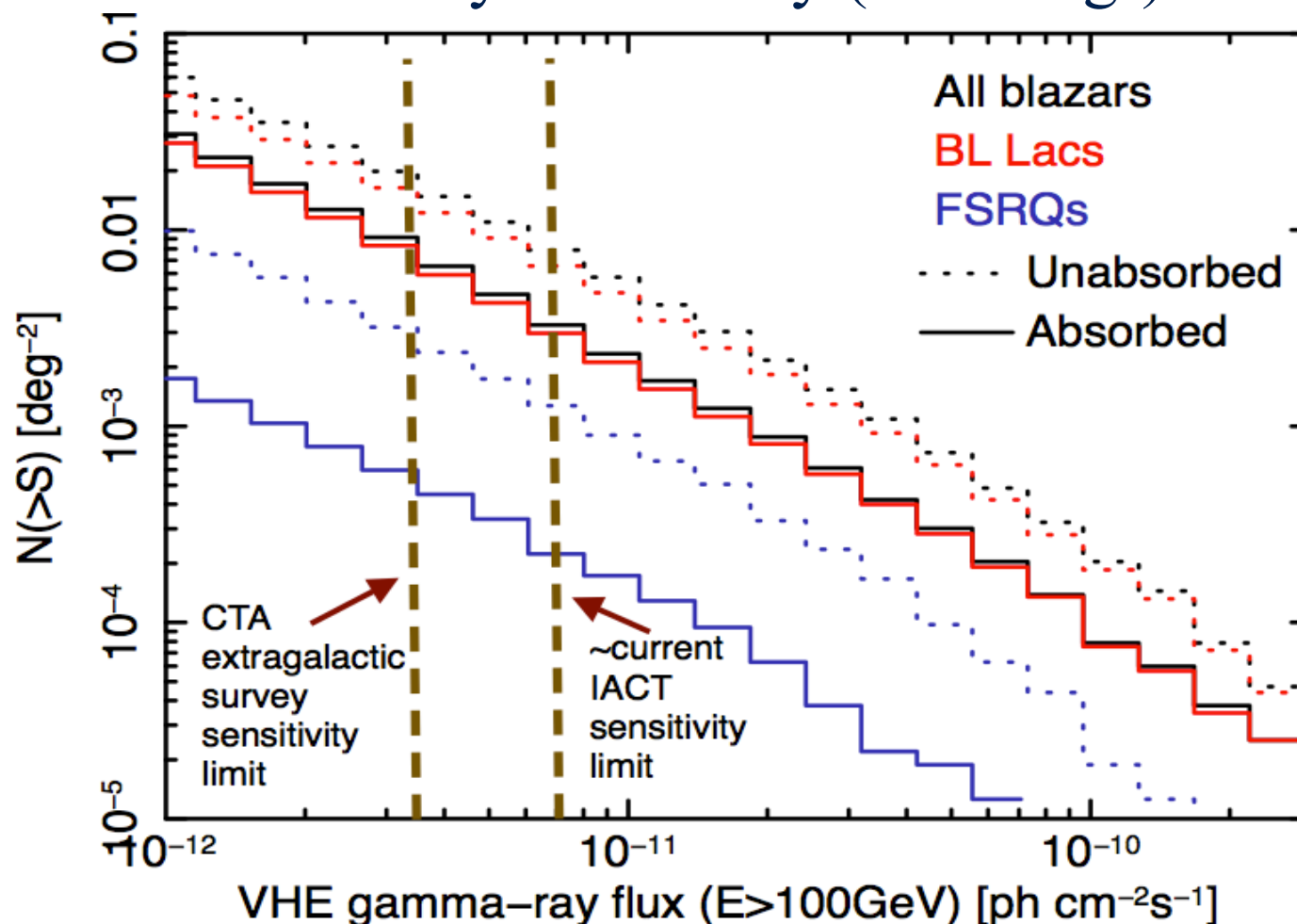
Unbiased survey for 1/4 sky ( $\sim 10^4$  deg $^2$ ) to flux limit  $\sim 5$ mCrab



# Extragalactic Survey



Unbiased survey for 1/4 sky ( $\sim 10^4 \text{ deg}^2$ ) to flux limit  $\sim 5 \text{ mCrab}$



Expect:

- 30-150 blazars (current total  $\sim 60$ )
- > clarify AGN  $\gamma$ LF, exgal.  $\gamma$  background
- new extreme blazars

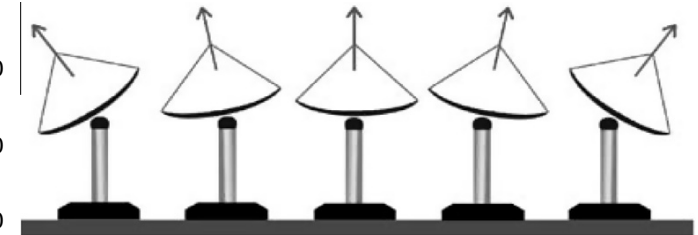
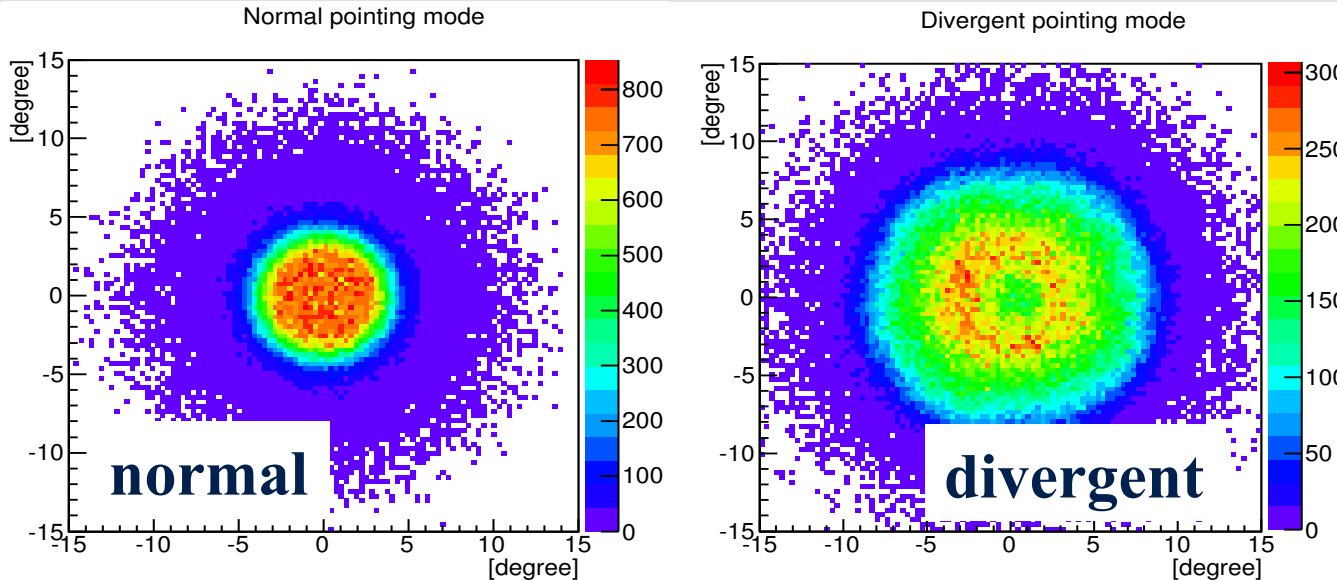
Discover new phenomena?

c.f. HESS unID

Fermi Bubble



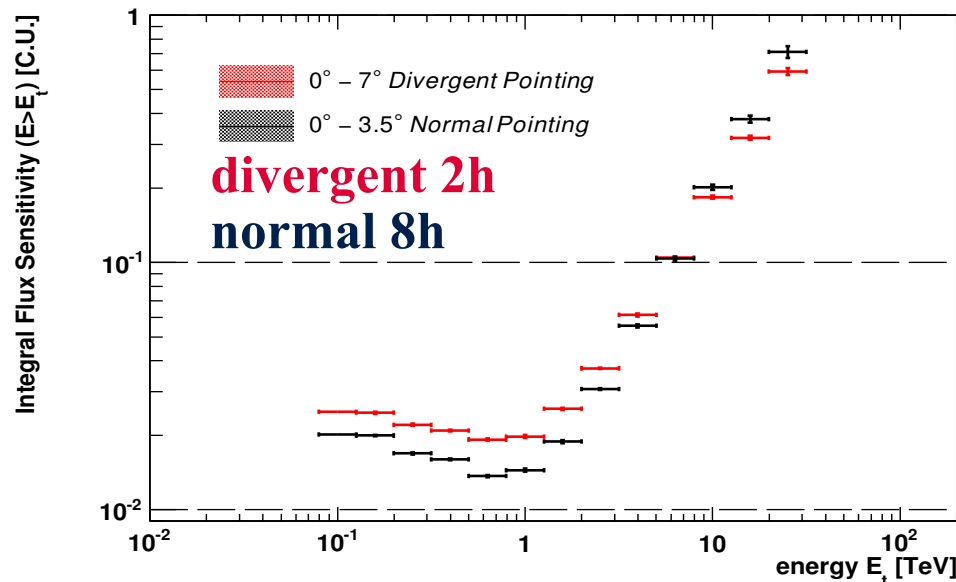
# Extragalactic Survey: Divergent Pointing?



Szanecki+ 2015

Wider FoV at  
expense of angular/  
energy resolution

Gérard+  
2015



Preliminary studies:  
point source survey  
efficiency comparable?  
-> possibly interesting  
for commensal unbiased  
transient survey

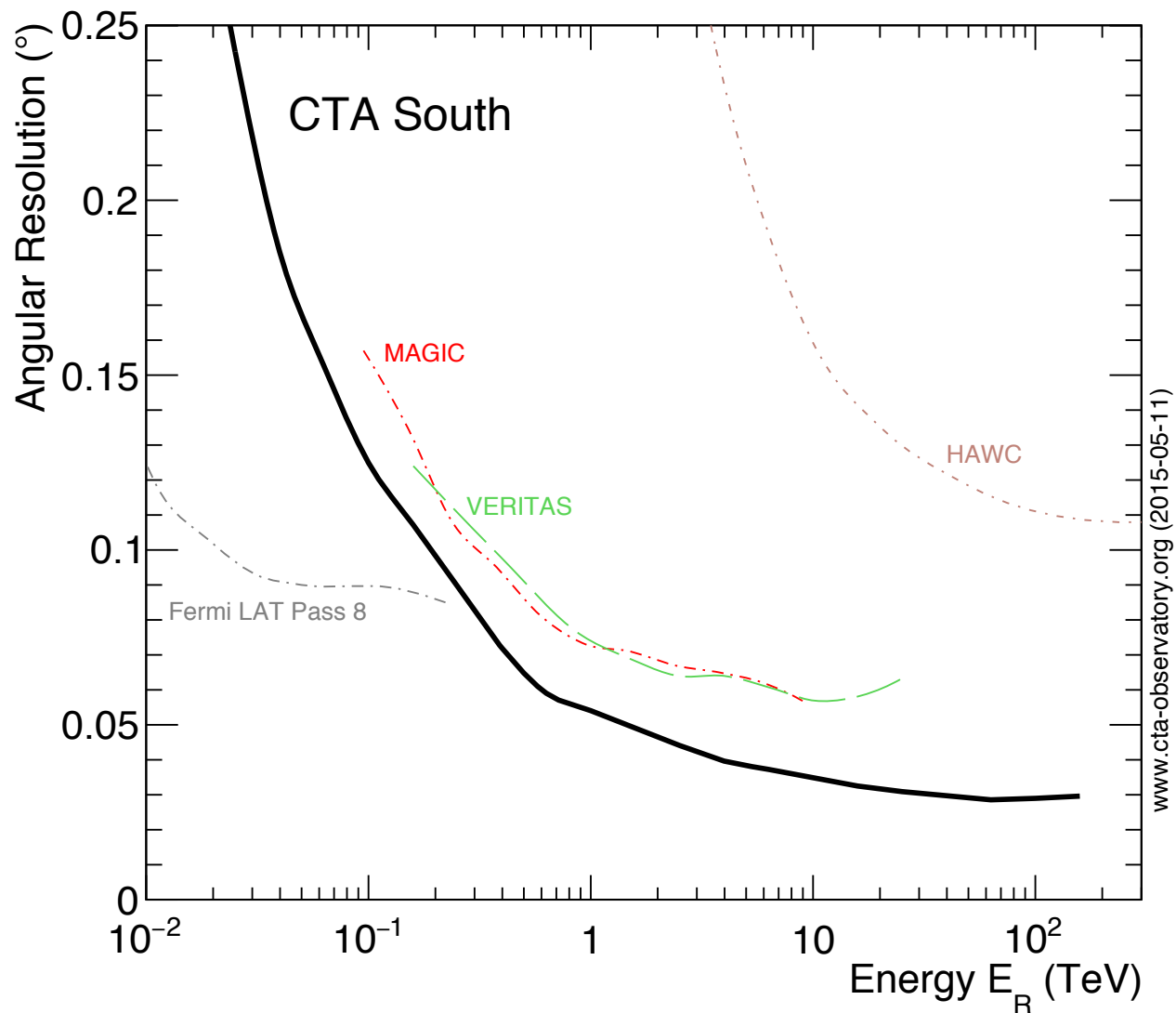
# Summary: CTA



- 
- **New ground-based gamma-ray observatory**
    - Open to the community
    - All-sky, high-sensitivity, wide-band (20 GeV - 300 TeV)
    - Rapid follow-up + rapid alerts for transients
    - Strong multi-wavelength/messenger synergy with concurrent projects
  - **Powerful tool for black hole astrophysics**
    - New perspectives on physics of AGN, GRBs,  $\mu$ quasars, etc
    - New approaches to observational cosmology, test of fundamental physics, etc
    - High-quality spectra + light curves for individual objects
    - Large survey programs for population studies, new discovery space

**Backup slides**

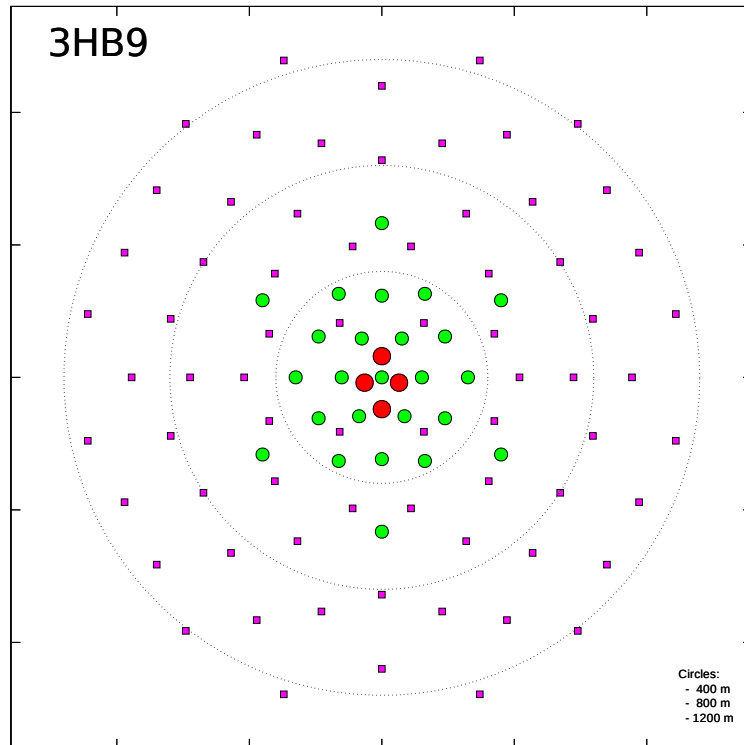
# CTA Angular Resolution



www.cta-observatory.org (2015-05-11)

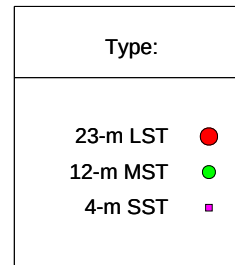


# Baseline Array Layout



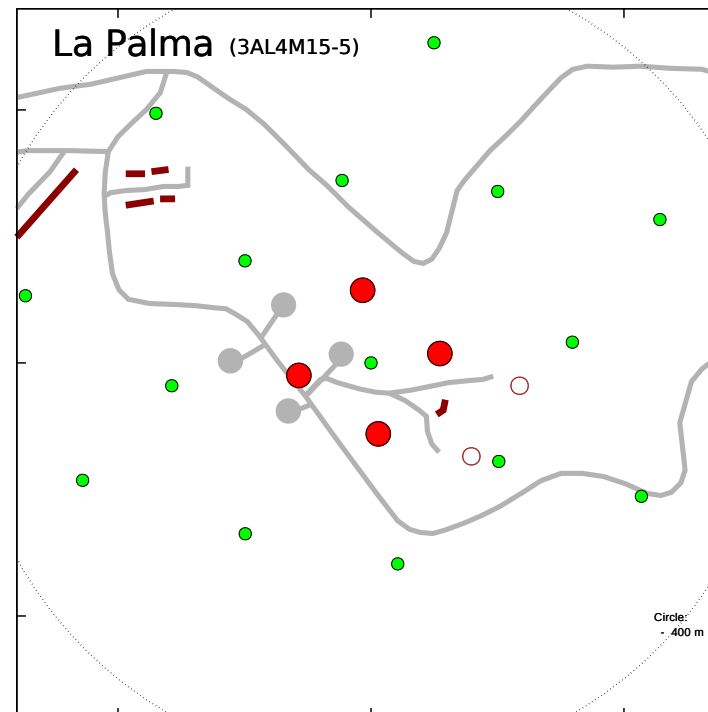
4 LSTs, 25 MSTs, 70 SSTs

**South:**  
 4 LST  
 25 MST  
 70 SST

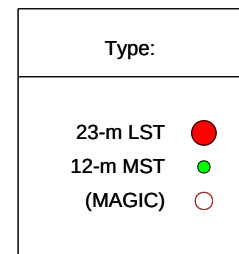


**North:**  
 4 LST  
 15 MST

1000 m



4 LSTs, 15 MSTs



North (x) is up  
 West (y) is left

250 m

from W. Hofmann

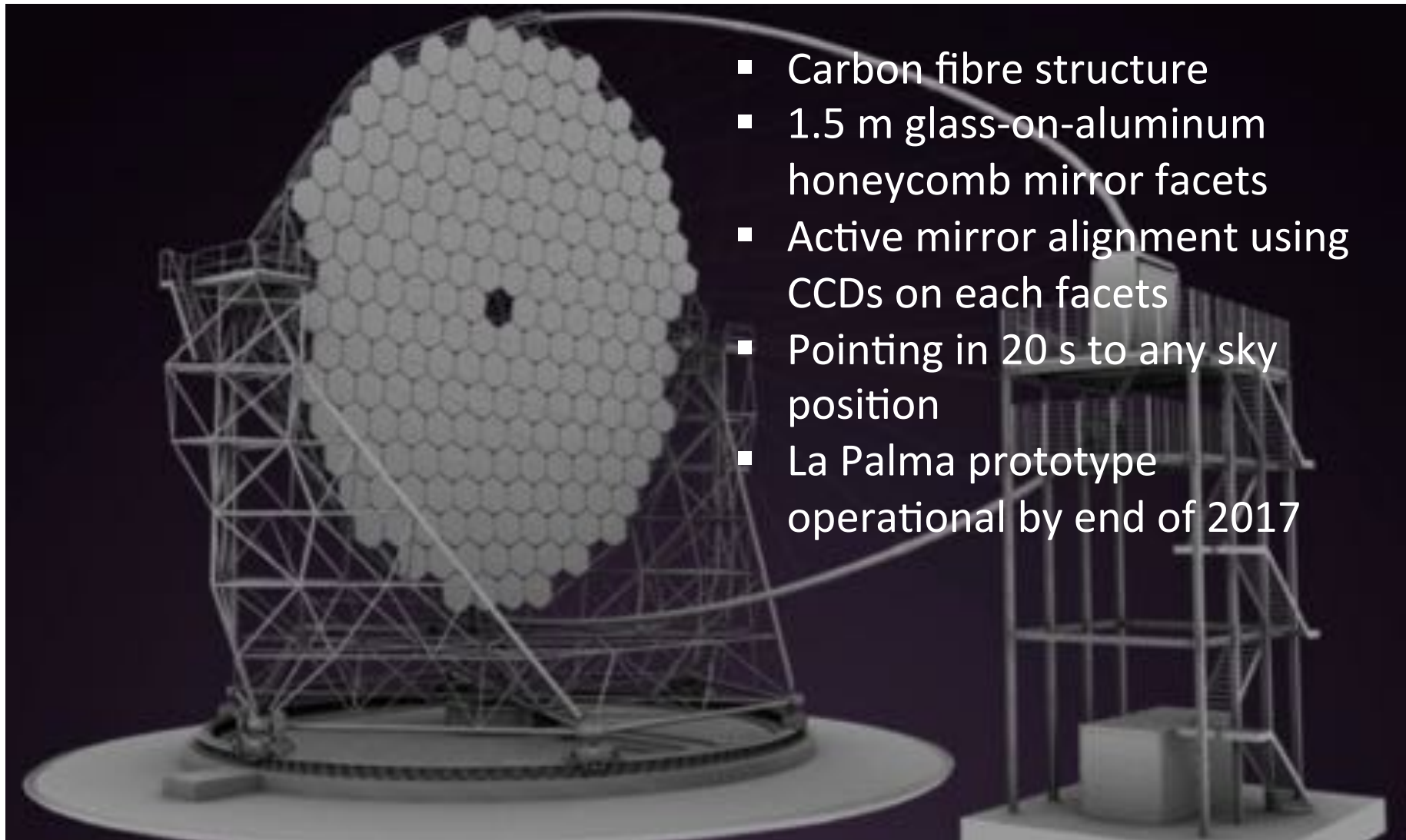
# Telescope Characteristics



| Telescope  | Large                    | Medium            |                      | Small             |                      |                      |
|--|--------------------------|-------------------|----------------------|-------------------|----------------------|----------------------|
|  | LST                      | MST               | SCT                  | SST-1M            | ASTRI SST-2M         | GCT SST-2M           |
| Number North array                                 | 4                        | 15                | TBD                  | 0                 |                      |                      |
| Number South array                                 | 4                        | 25                | TBD                  | 70                |                      |                      |
| <b>Optics</b>                                      |                          |                   |                      |                   |                      |                      |
| Optics layout                                      | Parabolic mirror         | Davies-Cotton     | Schwarzschild-Couder | Davies-Cotton     | Schwarzschild-Couder | Schwarzschild-Couder |
| Primary mirror diameter (m)                        | 23                       | 13.8              | 9.7                  | 4                 | 4.3                  | 4                    |
| Secondary mirror diameter (m)                      | –                        | –                 | 5.4                  | –                 | 1.8                  | 2                    |
| Eff. mirror area after shadowing (m <sup>2</sup> ) | 368                      | 88                | 40                   | 7.4               | 6                    | 6                    |
| Focal length (m)                                   | 28                       | 16                | 5.6                  | 5.6               | 2.15                 | 2.28                 |
| <b>Focal plane instrumentation</b>                 |                          |                   |                      |                   |                      |                      |
| Photo sensor                                       | PMT                      | PMT               | silicon              | silicon           | silicon              | silicon              |
| Pixel size (degr.), shape                          | 0.10, hex.               | 0.18, hex.        | 0.07, square         | 0.24, hex.        | 0.17, square         | 0.15-0.2, square     |
| Field of view (degr.)                              | 4.5                      | 7.7/8.0           | 8.0                  | 9.1               | 9.6                  | 8.5 - 9.2            |
| Number of pixels                                   | 1855                     | 1764/1855         | 11328                | 1296              | 1984                 | 2048                 |
| Signal sampling rate                               | GHz                      | 250 MHz / GHz     | GHz                  | 250 MHz           | S&H                  | GHz                  |
| <b>Structure</b>                                   |                          |                   |                      |                   |                      |                      |
| Mount  | alz-az, on circular rail | alt-az positioner | alt-az positioner    | alt-az positioner | alt-az positioner    | alt-az positioner    |
| Structural material                                | CFRP / steel             | steel             | steel                | steel             | steel                | steel                |
| Weight (full telescope, tons)                      | 100                      | 85                | ~85                  | 9                 | 15                   | 8                    |
| Max. time for repositioning (s)                    | 20                       | 90                | 90                   | 60                | 80                   | 60                   |

from W. Hofmann

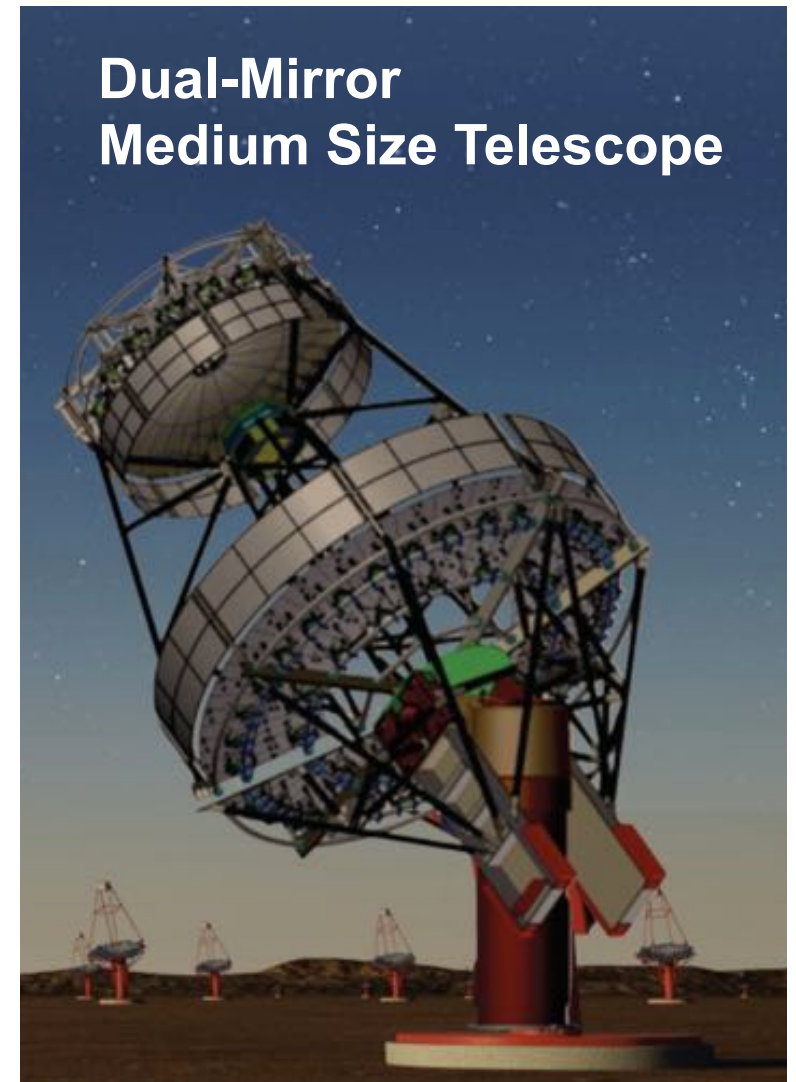
# Large Sized Telescope



- Carbon fibre structure
- 1.5 m glass-on-aluminum honeycomb mirror facets
- Active mirror alignment using CCDs on each facets
- Pointing in 20 s to any sky position
- La Palma prototype operational by end of 2017

from W. Hofmann

# Middle Sized Telescopes



adapted from W. Hofmann



# Small Sized Telescopes



**Meudon**



**Sicily**



**Kraków**



**adapted from W. Hofmann**